

## Atmospheric Results from the Mars Exploration Rovers

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Although at first glance, the Mars Exploration Rover (MER) payload may be perceived as primarily suited to geological investigation, it is in fact quite well-suited to carry out a robust and dynamic program of atmospheric monitoring and characterization with a particular emphasis on the planetary boundary layer. More to the point, it has been doing so at both the Gusev and Meridiani locations for more than one martian year. Ongoing atmospheric observations include (1) periodic thermal infrared spectra of the Martian sky by the Miniature Thermal Emission Spectrometer (Mini-TES). The actual sequences consist of both standard 200-second integrations and long “stares” of up to (almost) an hour. These data are highly diagnostic of vertical thermal structure (from 10 meters to 3 kilometers) and aerosol optical depths. (2) direct solar imaging using the Panoramic Camera (Pancam) and 440/880 nm + neutral density (ND5) filters, providing accurate measurement of visible optical depths. (3) near-sun and “sky-arc” sequences using the full suite of geological filters, intended to capture the forward-diffraction peak and the phase function characteristics of the aerosol particles. (4) carbon dioxide (15 micrometer band) profiling of the Mini-TES surface observations, providing an average near-surface (1 m) air temperature. The above activities have been (and will continue to be) used to characterize diurnal and secular temporal trends and to examine the spatial variability of such trends. In addition, serendipity has provided the unique opportunities of watching the decay of a moderate dust storm from two widely-separated sites as well as of multiple simultaneous orbiter-rover observing “campaigns.” The latter includes thus far the Mars Express and Mars Global Surveyor over-flights. During our presentation, we will briefly summarize the atmospheric results obtained and analyzed through the end of the first 700 sols (days) of operations, the unique contributions/capabilities of each instrument, and the synergy which comes from combining the two, e.g., visible-to-infrared optical ratio.