

Nuclear Spin Temperature of Methane in Comet C/2001 Q4

HIDEYO KAWAKITA¹, JUN-ICHI WATANABE², REIKO FURUSHO³, TETSUHARU FUSE⁴ and D.ANIEL C. BOICE⁵

¹Gunma Astronomical Observatory ²National Astronomical Observatory of Japan ³Waseda University ⁴Subaru Telescope, National Astronomical Observaotry of Japan ⁵Southwest Research Institute

Comets are primordial icy bodies in the Solar System. Formation or condensation conditions of cometary molecules remain in various properties of the molecules, e.g., chemical compositions. Here we concentrate on abundance ratio of nuclear spin isomers of cometary molecules. Cold nuclear spin temperatures of cometary water and ammonia have been reported so far. However, there are no reliable reports on a nuclear spin temperature of methane in comets.

Methane has four identical hydrogen nuclei and its rotational energy levels are classified into three nuclear spin states; *A*, *E*, and *F*. We observed comet C/2001 Q4 (NEAT) by the 10-meter Subaru telescope with the IRCS (Infrared Camera and Spectrograph) to determine the nuclear spin temperature of methane on 2004 May 28 UT. High-dispersion spectroscopic observations (R=10000) were performed by using the cross-dispersed echelle mode of the IRCS. R-branch emission lines (R0 – R6) of the v_3 vibrational band of methane were clearly detected. Cold nuclear spin temperature of methane has been derived to be 33 K at the first time in the comet.

Keywords: Comets; Nuclear spin temperature; Methane.

References

- [1] H. Kawakita et al., Astrophys. Journal 601, 1152 (2004).
- [2] N. Dello Russo et al., Astrophys. Journal, in press (2005).
- [3] D. Bockelee-Morvan et al., in COMETS II, in press (2005).
- [4] J. Crovisier, Faraday Discuss. 109, 437 (1998).